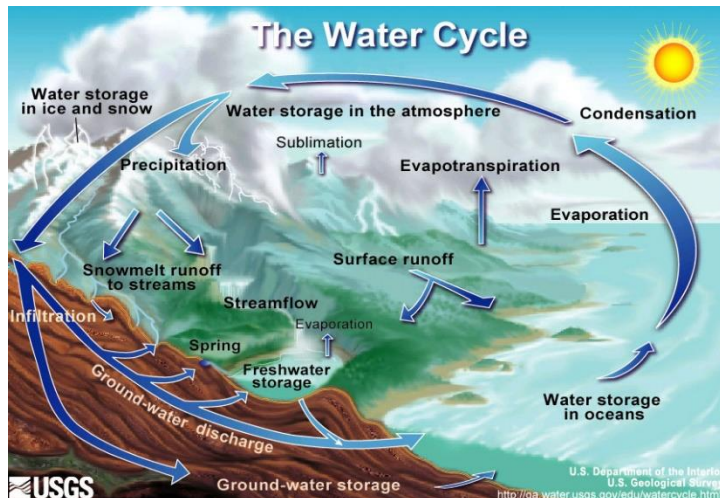
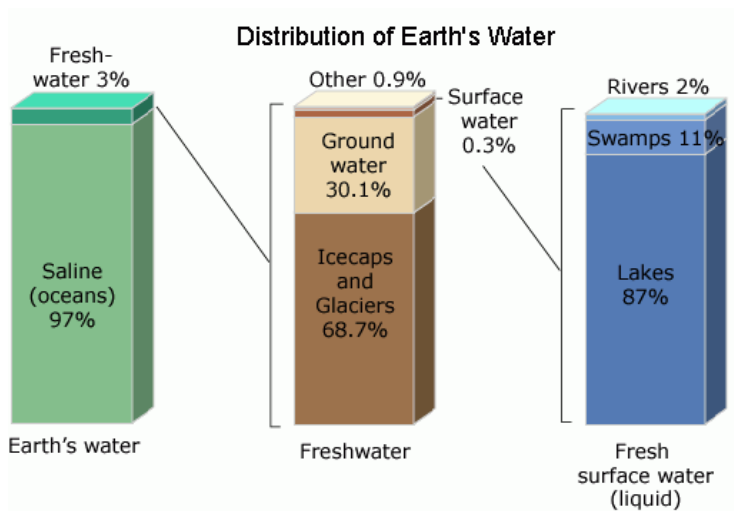


Hydrological Cycle

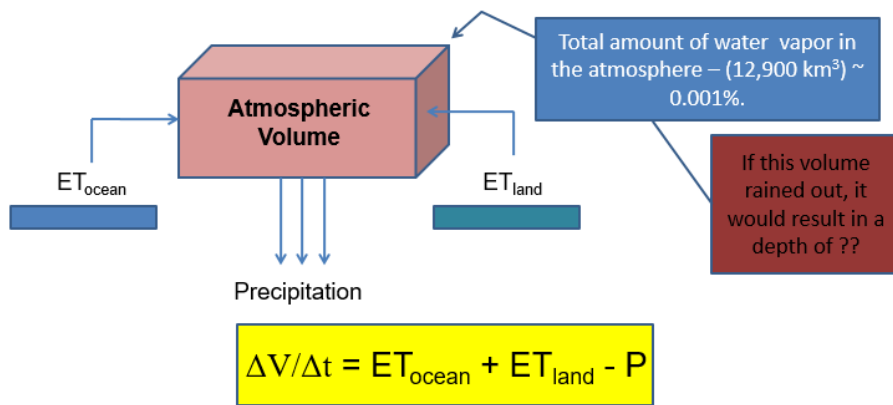


- More than 70% of the Earth is covered by water.
- The **hydrosphere** contains about $1.36 \times 10^9 \text{ km}^3$ of water
- Only substance (on Earth) that can exist in all states
- Maximum density in liquid form (not solid).
- Conducts heat & has a high **specific heat**.

Distribution



- Oceans make up ~ 70% of the Earth's surface.
- Solar radiation powers the global evaporation process.
- 86% of the Earth's evaporation occurs over the oceans; 14% occurs over land
- Of the total amount evaporated, P returns > 80% to the oceans and < 20% to land.
- Runoff sends 7% of the land based P back to the ocean to balance the processes of ET and P.



Residence times

- Amount of time that water 'stays' in a storage component can be calculated by dividing the **total volume (m³) by the flux rate (m³.s⁻¹)**.
- Water is renewed in rivers once every 16 days. Water in the atmosphere is completely replaced every 8 days, much longer in lakes, glaciers etc.

Stefan-Boltzmann Law

- Perfect emitter
 - o $E = \sigma T^4$
- Relationship says that a small increase in temperature results in a large amount of radiation being emitted.

Wein's Law

- There is an inverse relationship between the wavelength of the peak emission of a blackbody and its temperature, wavelength shorter as temperature increases
- $\lambda_m T = 2877 \mu m K$
- Light from the sun: temp ~ 6000K, $\lambda \sim 0.5 \mu m$
- Light from a campfire: temp ~ 1500K, $\lambda \sim 2 \mu m$

Units

- Joule – basic unit of energy
 - o 1 Joule work done to move 1 Newton by 1 metre
 - o (lift a small apple by one metre)
- Power – energy per time
 - o 1 Watt = 1 Joule per second
- Flux – power per area

Radiation Balance

$$S = S_o \left(\frac{r_o}{D} \right)^2$$

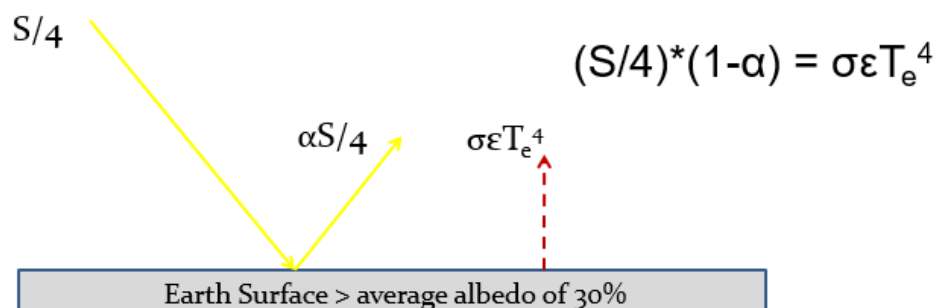
- S is the radiation at Earth, S_o is the radiation leaving the sun, r_o is the radius of the sun, D is the distance from the earth to the sun

Albedo

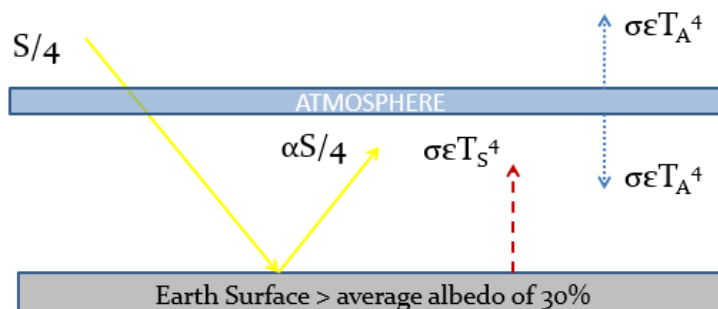
- Albedo = reflected solar radiation/incoming solar radiation
- On Earth, albedo = 30%
- Stefan Boltzmann relationship
- $E = \sigma \epsilon T_e^4$
 - o E is the energy flux in W/m²
 - o e is the emissivity of the surface, 1 for a blackbody
 - o σ is constant
 - o T is temp in Kelvin

Global radiation balance

Energy absorbed by Earth = Energy emitted by Earth



Energy absorbed by Earth = Energy emitted by Earth



This is the **radiation balance** at the Earth's surface

$$(S/4)*(1-\alpha) + \sigma \epsilon T_A^4 = \sigma \epsilon T_s^4$$

Seasonal Variation

- Changes in the position of the Earth's axis cause the location of the Sun to wander 47° across our skies.
- These changes have a direct effect on the **intensity** of solar radiation incident upon the surface.

Energy Distribution